

# Surface Energy Budget

Please read Bonan Chapter 13

## Energy Budget Concept

- For any system,  

$$(\text{Energy in}) - (\text{Energy out}) = (\text{Change in energy})$$
- For the land surface,
  - Energy in = ?
  - Energy Out = ?
  - Change in energy = ?

## Energy Budget Concept

- For the land surface,
  - Energy in = **Radiation**
  - Energy Out = **Radiation + Turbulent fluxes of “sensible” and “latent” heat**
  - Change in energy = **changes in temperature of soil, plants, water, and air**

## Surface Radiation Budget

- **Shortwave**
  - Down (solar constant, seasonal and diurnal geometry, atmospheric attenuation, clouds and aerosol)
  - Up (albedo)
- **Longwave**
  - Down (emission from atmosphere depends on temperature profile, water vapor, clouds)
  - Up (surface temperature, emissivity)

## Surface Radiation Budget

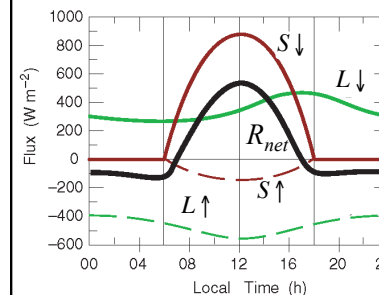
$$R_{net} = SW \downarrow - SW \uparrow + LW \downarrow - LW \uparrow$$

$$= SW \downarrow (1 - albedo) + LW \downarrow (1 - \epsilon) + \epsilon \sigma T_s^4$$

- **Shortwave**
  - Down (solar constant \* cosZ \* transmissivity<sup>airmass</sup>)
  - Up (albedo \* SW down)
- **Longwave**
  - Down (complicated! Weighted average of  $\sigma T_a^4$ )
  - Up ( $\epsilon \sigma T_s^4$ )

## Radiation Budget Components

Diurnal Cycle

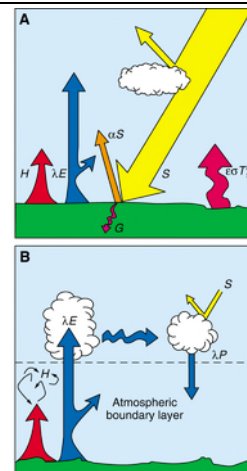


- Net solar follows cos(z)
- LW fluxes much less variable ( $\epsilon \sigma T^4$ )
  - LW up follows surface T as it warms through day
  - LW down changes little
  - LW net opposes  $SW_{net}$
- $R_{net}$  positive during day, slightly negative at night

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## Land Surface Energy Budget

- Very little of the energy gained by net radiation is stored in the ground (**G**)
- Most is emitted as LW IR and turbulent fluxes of sensible (**H**) and latent heat (**LE**)
- Latent energy is then **released into atmosphere** when vapor condenses



**Fig. 1.** Interactions between the land surface and the atmosphere that have direct impacts on the physical climate system. **(A)** Surface radiation budget. **(B)** Effect of heat fluxes on the atmosphere.

## Surface Energy Budget

**Storage change = Energy in – energy out**

$$\rho c \frac{\Delta T}{\Delta t} \Delta z = (S \downarrow - S \uparrow + L \downarrow - L \uparrow) - H + \lambda E = G$$

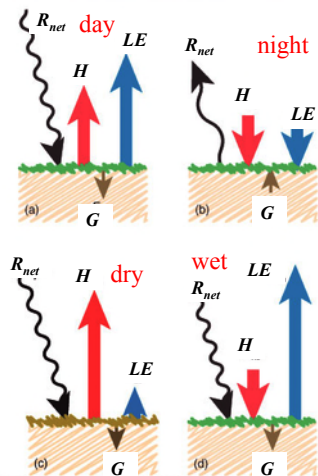
$$R_{net} = (S \downarrow - S \uparrow) + (L \downarrow - L \uparrow) = H + \lambda E + G$$

↑ net SW     ↑ net LW     ↑ sensible heat flux     ↑ Latent heat of vaporization     ↑ evap     ↑ ground heat flux

**Role of the land surface:**

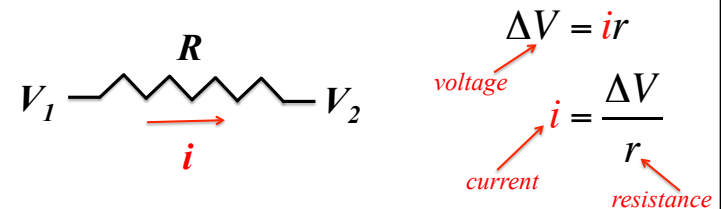
**Partition of net radiation into turbulent fluxes & storage**

## Surface Energy Budgets



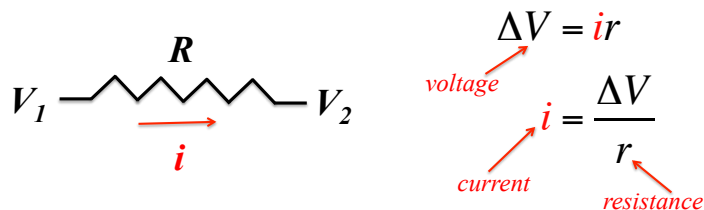
- $R_{\text{net}} = H + LE + G$   
 $\sim H + LE$
- Daytime turbulent fluxes upward
- Night: turbulent fluxes downward (dew or frost!)
- Dry surfaces  
 $R_{\text{net}} \sim H$
- Wet surfaces  
 $R_{\text{net}} \sim LE$

## Remember Ohm's Law?



- **Flow of current** through a resistor is ratio of difference in **potential** to **resistance** (this is just a “definition of resistance”)
- This is another form of our familiar concept of stuff flowing from high concentration to low concentration (like “Fickian **Diffusion**”)

## Heat Fluxes ~ Currents



- **Sensible heat flux**
  - Driving potential is a difference in temperature
  - **H** is proportional to  $\Delta T$
- **Latent heat flux**
  - Driving potential is a difference in vapor pressure
  - **LE** is proportional to  $\Delta e$

## Sensible Heat Flux

- Driving potential is a difference in temperature
- **H** is proportional to  $\Delta T$

$$H = \rho c_p \frac{\Delta T}{r} = \rho c_p \frac{T_s - T_a}{r}$$

Watts  $\text{m}^{-2}$       density  $\text{kg m}^{-3}$       heat capacity  $1004 \text{ J K}^{-1} \text{ kg}^{-1}$       aerodynamic resistance  $\text{s m}^{-1}$

$$\frac{(\text{kg m}^{-3})(\text{J K}^{-1} \text{ kg}^{-1})(\text{K})}{(\text{s m}^{-1})} = \frac{(\text{kg m}^{-3})(\text{J K}^{-1} \text{ kg}^{-1})(\text{K})}{(\text{s m}^{-1})} = \frac{\text{J s}^{-1}}{\text{m}^2} = \text{W m}^{-2}$$

## Water Vapor Pressure

- Molecules in an air parcel **all contribute to pressure**
- Each subset of molecules (e.g., N<sub>2</sub>, O<sub>2</sub>, H<sub>2</sub>O) exerts a **partial pressure**
- The **VAPOR PRESSURE ( $e$ )**, is the pressure exerted by water vapor molecules in the air

## Latent Heat Flux

- Driving potential is a difference in water vapor pressure
- **LE** is proportional to  $\Delta e$

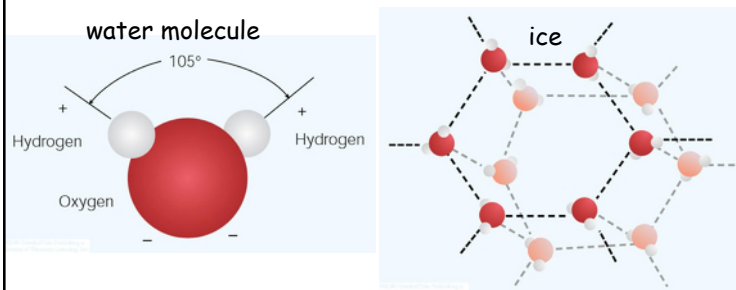
$$LE = \frac{\rho c_p}{\gamma} \frac{\Delta e}{r} = \frac{\rho c_p}{\gamma} \frac{e_s - e_a}{r}$$

Watts m<sup>-2</sup>      "Psychrometric constant"      aerodynamic resistance s m<sup>-1</sup>

$$\gamma = (C_p P) / (0.622 \lambda)$$

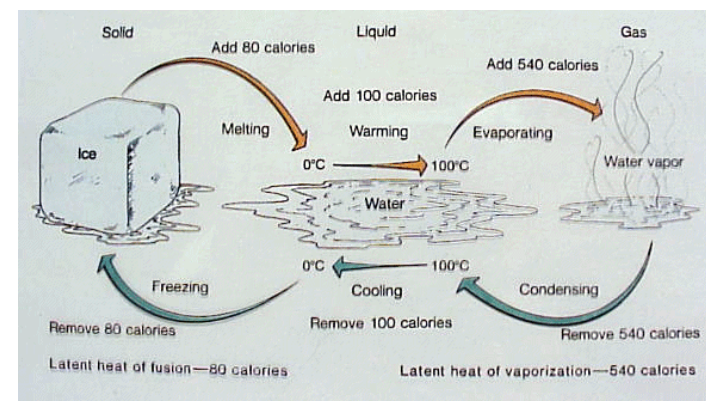
$$66.5 \text{ Pa } ^\circ\text{C}^{-1}$$

## Molecular Structure of Water



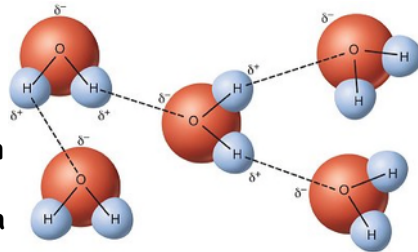
Water's unique molecular structure and hydrogen bonds enable all 3 phases to exist in Earth's environments

## "Latent" (hidden) Energy associated with phase changes



## Why does it take so much energy to evaporate water?

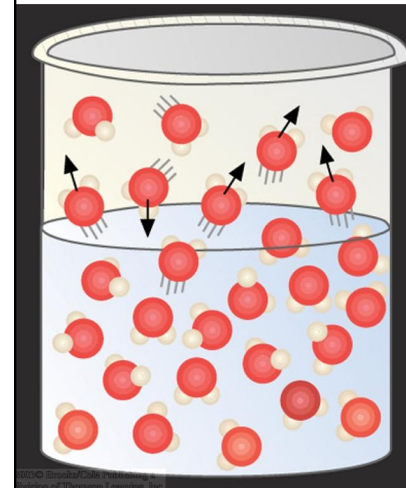
- In the liquid state, adjacent water molecules **attract** one another
- This same hydrogen bond accounts for **surface tension** on a free water surface



*"plus" charge on hydrogen in one water molecule attracts the "minus" charge on a neighbor's oxygen*

**Evaporation must break these hydrogen bonds**

## Water Vapor Saturation



- Water molecules **move** between the liquid and gas phases
- When the rate of water molecules entering the liquid equals the rate leaving the liquid, we have **equilibrium**
- The air is said to be **saturated** with water vapor at this point

## Clausius-Clapeyron Eqn

(see Monteith & Unsworth, pp 11-13)

- From Second Law of Thermodynamics

$$\frac{de_s}{dT} = \frac{L}{T(\alpha_2 - \alpha_1)}$$

*change in saturation vapor pressure* (pointing to  $de_s$ )  
*temperature* (pointing to  $dT$ )  
*Latent heat of vaporization* (pointing to  $L$ )  
*specific volumes of liq & vapor* (pointing to  $\alpha_2 - \alpha_1$ )

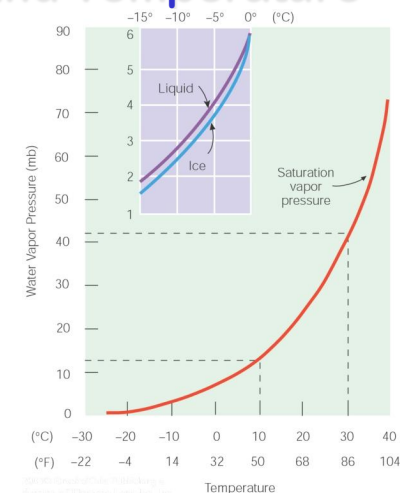
- Approximate but very accurate 0° to 35° C

$$e_s(T) = e_s(T^*) \exp\{A(T - T^*)/(T - T')\}$$

where  $A = 17.27$ ,  $T^* = 273$  K ( $e_s(T^*) = 0.611$  kPa), and  $T' = 36$  K.

## Saturation and Temperature

- The **saturation vapor pressure** of water increases with temperature
  - At higher T, faster water molecules in liquid escape more frequently causing **equilibrium** water vapor concentration to rise
  - We sometimes say "**warmer air can hold more water**"
- There is also a vapor pressure of water over an ice surface
  - The **saturation vapor pressure above solid ice is less than above liquid water**



## Latent Heat Flux

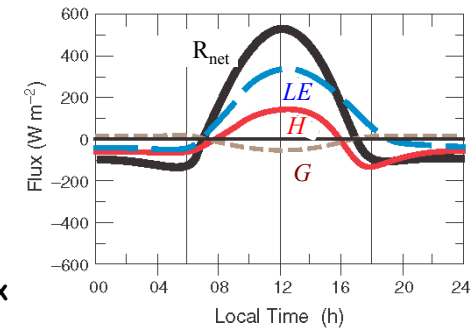
- Driven by difference in vapor pressure
- Over open water  $e_{\text{surface}} = e_{\text{sat}}(T_s)$
- Over vegetation, liquid water is evaporating inside tiny openings in leaves called **stomata** (singular “**stomate**”)
- Evapotranspiration = latent heat flux is driven by the **vapor pressure deficit**

$$\text{vpd} = (e_{\text{sat}}(T_s) - e_a)$$

$$LE = \frac{\rho c_p}{\gamma} \frac{e_s - e_a}{r} = \frac{\rho c_p}{\gamma} \frac{e_{\text{sat}}(T_s) - e_a}{r}$$

## Idealized Diurnal Cycle

- $R_{\text{net}}$  follows  $\cos(z)$  during day, negative at night (LW cooling)
- Downward turbulent fluxes at night
- Ground heat flux smaller: downward during day and up at night

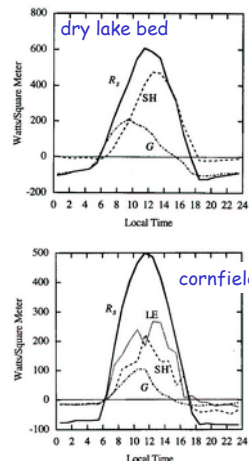


$$R_{\text{net}} = H + LE + G$$

$$\sim H + LE$$

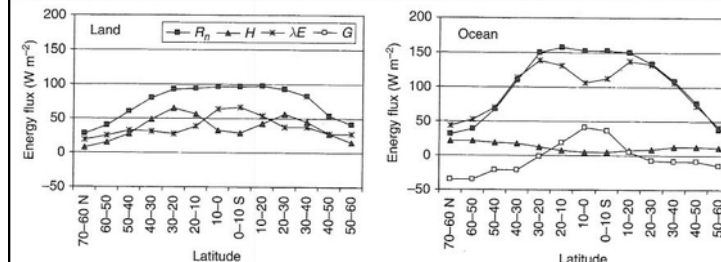
## Energy Budget Components

### Observed Diurnal Cycles



- Recall  $R_{\text{net}} = H + LE + G$
- $R_{\text{net}}$  follows  $\cos(z)$  during day
- $R_s < 0$  at night due to **LW cooling** at surface
- Over **desert**,  $LE = 0$  (dry), so  $R_{\text{net}}$  balanced early by  $G$ , later by  $H$
- Over **active vegetation**,  $R_{\text{net}}$  balanced by all three terms
  - $H + LE \gg G$
  - Note **dip in LE at mid-day** ... why?

## Global Variations



- Tropics  $\rightarrow$  pole gradients
- Lower albedo over oceans (higher  $R_{\text{net}}$ )
- $G \sim 0$  : **Storage negligible over land**
- **Water budget** in oceans and required **atmospheric transport**



## Radiation, Hydrology, & the Sfc Energy Budget

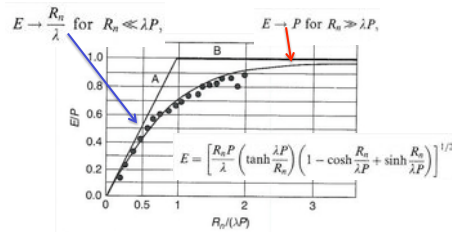
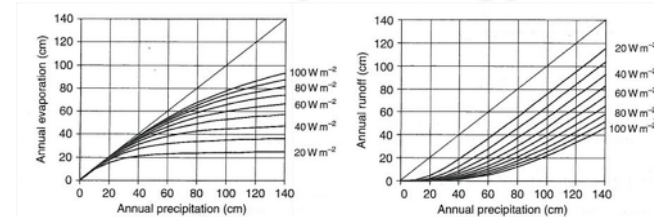


FIGURE 13.2. Relationship between the ratio of annual evaporation to precipitation ( $E/P$ ) and the ratio of net radiation to the amount of energy required to evaporate the annual precipitation ( $R_n/(\lambda P)$ ). Redrawn from Budyko (1974, p. 325).

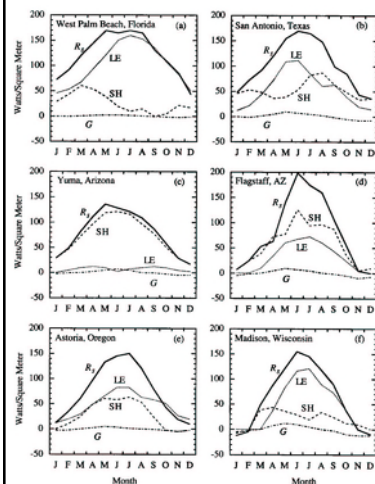
- **Wet places:** evaporation balances radiation
- **Dry places:** evaporation balances precipitation
- Smooth transition in between

## Annual Radiation and Hydrology



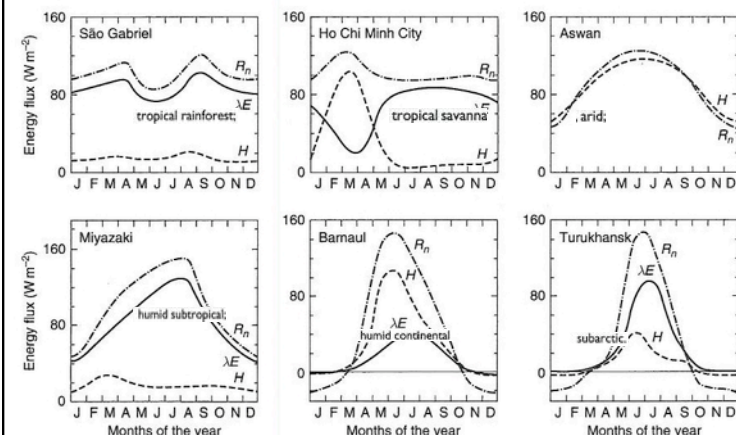
- Radiation favors evaporation over runoff
- Dry (sunny) places:  $LE \sim \text{Precipitation}$
- Wet (cloudy) places:  $\text{Runoff} \sim \text{Precipitation}$

## Seasonal Energy Budgets

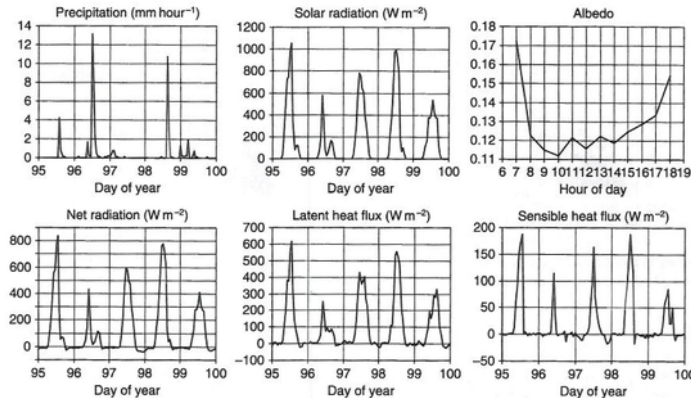


- Seasonal course of  $R_{net}$  due to **Sun-Earth geometry**
- **Moist** climates feature near balance of  $R_{net} \sim LE$
- **Dry** climates feature near balance of  $R_{net} \sim H$
- Others are intermediate
  - Spring vs fall in Texas
  - Summer (leaves) vs spring and fall in Wisc
- $(H, LE) \gg G$  everywhere

## Seasonal Energy Fluxes

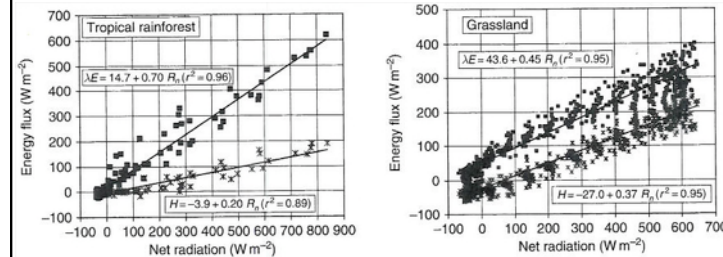


## Diurnal Variations



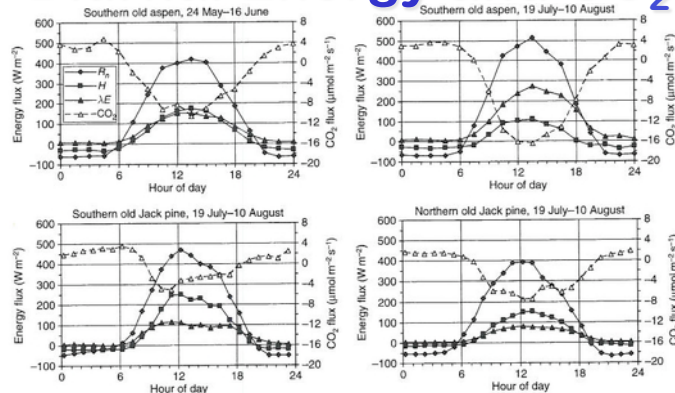
Tropical Forest: Rondonia, Brazil (10° S)

## Partition of Net Radiation



- Energy budget “closure”
- Forest “harvests” radiation to extract water from soil
- Grassland passes more energy back to atmosphere as sensible heat

## Diurnal Energy and CO<sub>2</sub>



- CO<sub>2</sub> flux is a mirror image of LE (stomatal control)
- Physiological differences: broadleaf vs needleleaf

## Partition of Net Radiation

$$R_{net} = (S \downarrow - S \uparrow) + (L \downarrow - L \uparrow) = H + \lambda E + G$$

*Ground heat flux*  
 $G = k(T_s - T_g) / \Delta z$

*Sensible flux driven by  $\Delta T$*

$$H = -\rho C_p \frac{(T_a - T_s)}{r_H}$$

*Latent flux driven by VPD*

$$\lambda E = -\frac{\rho C_p (e_a - e_s(T_s))}{\gamma r_w}$$

$\gamma = (C_p P) / (0.622 \lambda) \quad 66.5 \text{ Pa } ^\circ\text{C}^{-1}$   
*“Psychrometric constant”*



## Surface Energy Budget

Energy in = energy out + storage change

$$(1-r)S_{\downarrow} + \varepsilon L_{\downarrow} = \varepsilon \sigma (T_s + 273.15)^4 + H + \lambda E + G$$

$$(1-r)S_{\downarrow} + \varepsilon L_{\downarrow} = \varepsilon \sigma (T_s + 273.15)^4 - \rho C_p \frac{(T_a - T_s)}{r_H} - \frac{\rho C_p (e_a - e_s(T_s))}{\gamma} \frac{1}{r_W} + k \frac{(T_s - T_g)}{\Delta z}$$

- Can “solve” for surface temperature
- Physical properties: albedo, emissivity, heat capacity, soil conductivity & temperature
- “Resistances” are properties of the turbulence ... **depend sensitively on H!**

## Clausius-Clapeyron Eqn

(see Monteith & Unsworth, pp 11-13)

- **Approximate but very accurate 0° to 35° C**

$$e_s(T) = e_s(T^*) \exp\{A(T - T^*)/(T - T')\}$$

where  $A = 17.27$ ,  $T^* = 273$  K ( $e_s(T^*) = 0.611$  kPa), and  $T' = 36$  K.

- **Slope**  $s = \Delta(e_s)/\Delta T$  accurate from 0° to 40° C

$$s = \lambda M_w e_s(T) / (RT^2)$$

$$\lambda = L = 2.48 \text{ J kg}^{-1} \quad M_w (\text{mol wt water}) = 0.018 \text{ mol kg}^{-1}$$

$$R = 8.314 \text{ J mol}^{-1} \text{ K}^{-1} \text{ (“universal gas constant”)}$$

## Penman-Monteith Equation

“Thermodynamic” energy balance

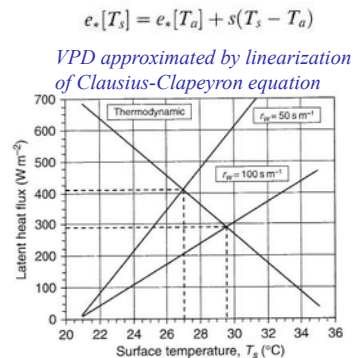
$$\lambda E = (R_n - G) - H = (R_n - G) + \rho C_p (T_a - T_s) / r_H$$

“Turbulent” energy balance

$$\lambda E = \frac{\rho C_p (e_s(T_s) + s(T_s - T_a) - e_a)}{\gamma}$$

Solve for surface temperature

$$T_s - T_a = (r_H / \rho C_p) (R_n - G - \lambda E)$$



## Solutions to P-M Equation

Latent heat flux

$$\lambda E = \frac{s(R_n - G) + \rho C_p (e_s(T_a) - e_a) / r_H}{s + \gamma(r_H / r_H)}$$

Sensible heat flux

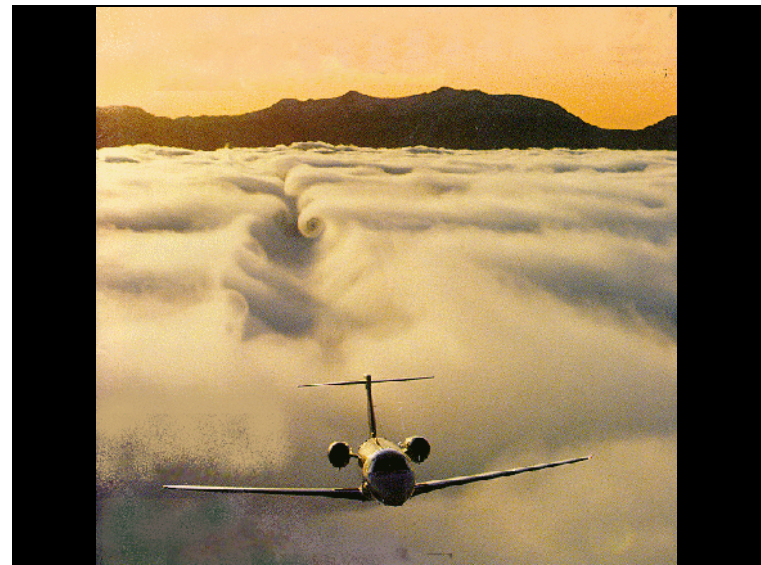
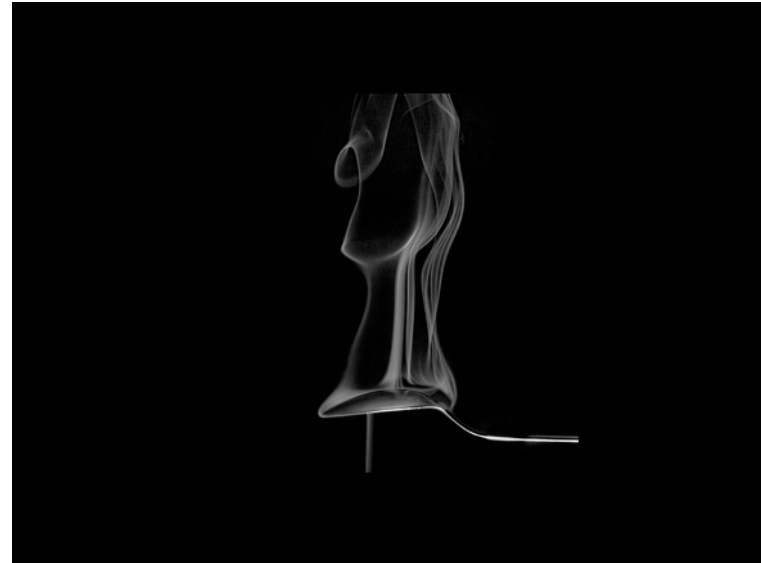
$$H = \frac{(R_n - G)\gamma^* - \rho C_p (e_s(T_a) - e_a) / r_H}{s + \gamma^*}$$

Surface temperature

$$T_s = T_a + \frac{(R_n - G)\gamma^* r_H / \rho C_p - (e_s(T_a) - e_a)}{s + \gamma^*}$$

# Turbulent Fluxes

Please read Bonan, Chapter 14



## Richardson's Rhyme

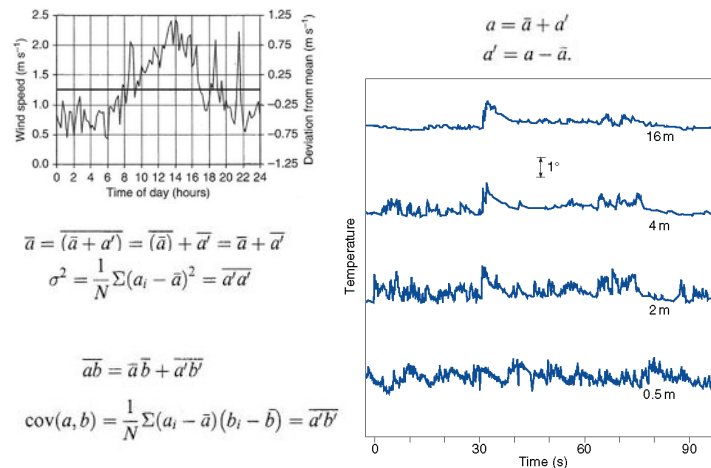
- “Big whorls have little whorls,  
Which feed on their velocity;  
And little whorls have lesser whorls,  
And so on to viscosity”  
– Lewis Richardson, *The supply of energy from and to Atmospheric Eddies* 1920
- “Great fleas have little fleas  
Upon their backs to bite 'em,  
And little fleas have lesser fleas,  
And so, ad infinitum”  
– Augustus De Morgan  
(19th century mathematician, parodying Jonathon Swift, 1733)

## Sonic Anemometer



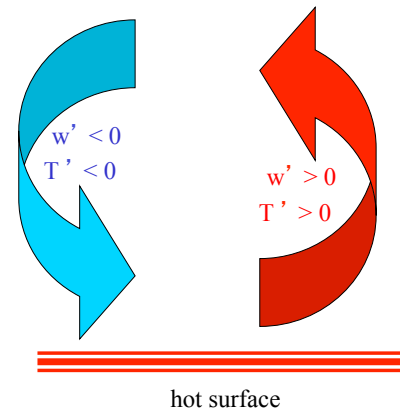
- Measures elapsed time for sound pulses to cross air in 3D
- Speed of sound is a known function of temperature
- Relative motion determined accurately in 3D
- Very fast instrument response time

## Time Series of Turbulence



## Turbulent Heat Flux

$$w \equiv \bar{w} + w' \quad T \equiv \bar{T} + T'$$



- Imagine a turbulent eddy over a hot surface
- Updrafts are systematically warmer than downdrafts
- Updraft:  
 $\overline{w'T'} > 0$
- Downdraft:  
 $\overline{w'T'} > 0$

## Sensible Heat Flux (Reynolds' Averaging)

Upward sensible heat flux  $H = c_p \rho w T$

Heat capacity at constant pressure  
 $1004 \text{ J K}^{-1} \text{ kg}^{-1}$

density of air  $\rho/RT$   
( $\sim 1.2 \text{ kg m}^{-3}$ )

Vertical velocity  
( $\text{m s}^{-1}$ )

Temp  
(K)

$w = \bar{w} + w'$ ,  $T = \bar{T} + T'$

Total time mean turbulent fluctuation

$$\begin{aligned}\overline{wT} &= \overline{(\bar{w} + w')(\bar{T} + T')} \\ &= \overline{\bar{w}\bar{T} + \bar{w}T' + w'\bar{T} + w'T'} \\ &= \overline{\bar{w}\bar{T}} + \overline{\bar{w}T'} + \overline{w'\bar{T}} + \overline{w'T'} \\ &= \bar{w}\bar{T} + \cancel{\overline{w'T'}} + \cancel{\overline{w'\bar{T}}} + \overline{w'T'}\end{aligned}$$

$$\overline{wT} = \bar{w}\bar{T} + \overline{w'T'}$$

- Mean of a mean is the mean
- Mean of a prime is zero
- Mean of a product is **not** necessarily zero

## Variance, Covariance, Correlation

$$\bar{c} = \frac{1}{N} \sum_{i=1}^N c_i \quad \text{mean}$$

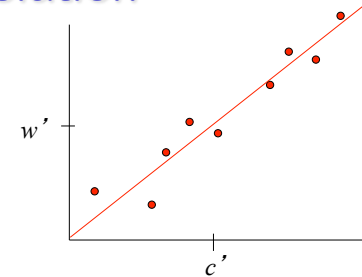
$$c_i' = c_i - \bar{c} \quad \text{perturbation}$$

$$\overline{c_i' c_i'} = \frac{1}{N} \sum_{i=1}^N (c_i - \bar{c})^2 = \sigma_c^2 \quad \text{variance}$$

$$\overline{w'^2} = \sigma_w^2 \quad \text{variance of } w$$

$$\overline{w' c'} = \frac{1}{N} \sum_{i=1}^N (c_i - \bar{c})(w_i - \bar{w}) = \text{cov}(w, c) \quad \text{covariance of } w \text{ and } c$$

$$\frac{\overline{w' c'}}{\sigma_w \sigma_c} = r \quad \text{Normalized covariance is the correlation coefficient}$$



## Turbulent Fluxes

### Sensible Heat Flux

$$\overline{wT} = \bar{w}\bar{T} + \overline{w'T'}$$

Total = mean + eddy

$$\overline{wq} = \bar{w}\bar{q} + \overline{w'q'}$$

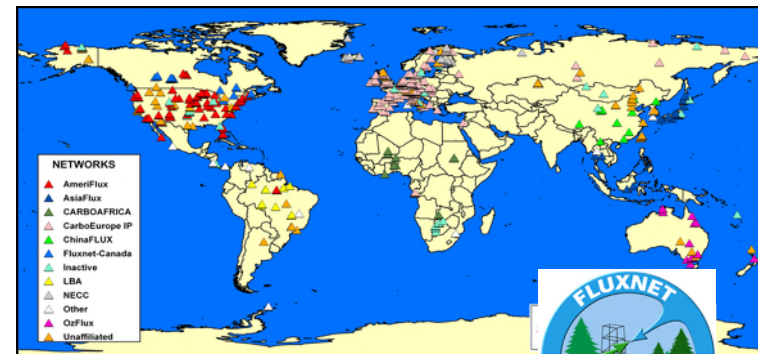
### Latent Heat Flux

Weird units:

$\text{K m s}^{-1}$

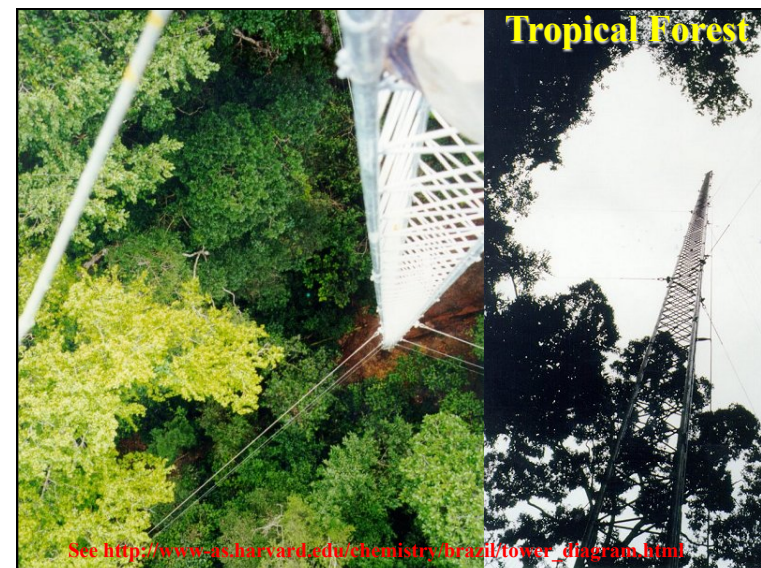
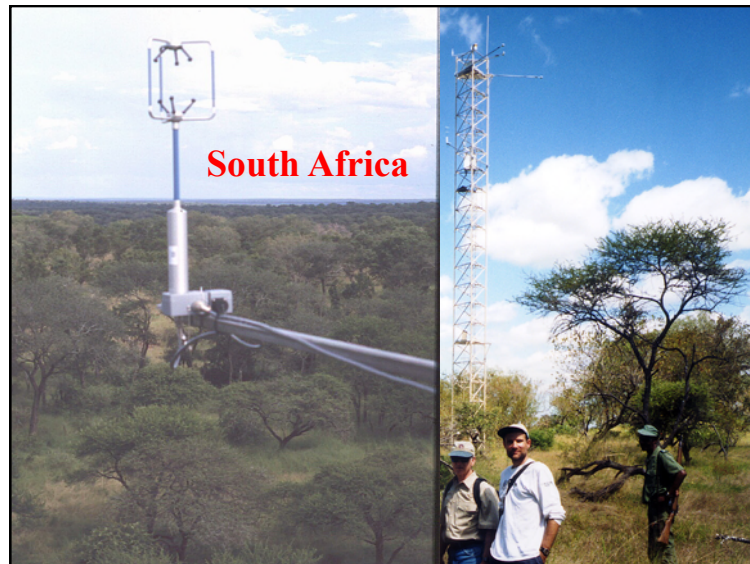
$\text{kg kg}^{-1} \text{ m s}^{-1}$

- Near the ground,  $\bar{w} \approx 0$
- “Eddy” terms dominate
- How can these fluxes be measured?



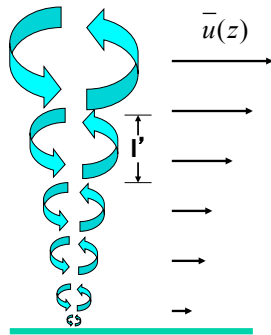
- **FluxNet:** An international “network of networks” > 550 sites
- 10 Hz measurements from many sites for > 5 yrs
- **H, LE, and NEE** of  $\text{CO}_2$  at most sites
- Data available online: <http://fluxnet.ornl.gov>







## Surface-Layer Mixing



- Turbulent eddies near the surface act to **mix atmospheric properties** ( $T, q, u$ ) and reduce vertical gradients
- Assume a **characteristic length scale  $l'$  for eddy mixing**, then

$$u = \bar{u}(z) + u'$$

$$u' = -l' \frac{\partial \bar{u}}{\partial z}$$

If eddies are **isotropic** (length and depth similar), then

$$|w'| \sim |u'|, \text{ so } w' \sim l' \frac{\partial \bar{u}}{\partial z}$$

## Surface Layer Stress

$$\begin{aligned} \tau_x &= -\rho \overline{w'u'} \\ &= -\rho \left( -l' \frac{\partial \bar{u}}{\partial z} \right) \left( -l' \frac{\partial \bar{u}}{\partial z} \right) = -\rho l'^2 \left| \frac{\partial \bar{u}}{\partial z} \right|^2 \\ &= K_m \frac{\partial \bar{u}}{\partial z}, \text{ where } K_m = \rho l'^2 \left| \frac{\partial \bar{u}}{\partial z} \right| \end{aligned}$$

Define  $u_* = \sqrt{\frac{\tau_x}{\rho}} = (\overline{u'w'})^{1/2}$  then

$$\frac{\tau_x}{\rho} = \frac{K_m}{\rho} \frac{\partial \bar{u}}{\partial z} = u_*^2$$

- Momentum flux** (surface stress) is proportional to the square of the product of the wind speed gradient (shear) and the turbulent length scale
- Define an “eddy viscosity” or “**eddy diffusivity**”  $K_m$  which is analogous to molecular diffusivity
- Define a **velocity scale  $u_*$**  for the turbulent eddies near the surface, called the **friction velocity**

## Surface Layer (cont' d)

$$\frac{\tau_x}{\rho} = \frac{K_m}{\rho} \frac{\partial \bar{u}}{\partial z} = u_*^2$$

$$K_m = \rho l'^2 \left| \frac{\partial \bar{u}}{\partial z} \right| = \rho k^2 z^2 \left| \frac{\partial \bar{u}}{\partial z} \right|$$

$$\frac{K_m}{\rho} \frac{\partial \bar{u}}{\partial z} = u_*^2 = \frac{\rho k^2 z^2 \left| \frac{\partial \bar{u}}{\partial z} \right|}{\rho} \frac{\partial \bar{u}}{\partial z}$$

$$\left( \frac{\partial \bar{u}}{\partial z} \right)^2 = \frac{u_*^2}{k^2 z^2}$$

$$\frac{\partial \bar{u}}{\partial z} = \frac{u_*}{kz}$$

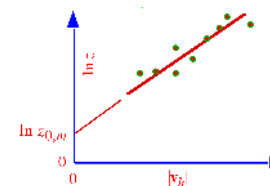
- Near the surface, **eddies are limited in size** by the proximity of the ground, so  $l'$  in  $K_m$  is  $l'(z)$
- Assume  $l' = kz$ , where  $k \sim 0.4$  is an empirical coefficient known as “**von Karman’s constant**”
- Leads to a characteristic relationship for variation of mean wind speed with height: the **log-wind profile**

## Log Wind Profile

$$\frac{\partial \bar{u}}{\partial z} = \frac{u_*}{kz} \text{ separate variables}$$

$$\int_{z_0}^z \frac{\partial \bar{u}}{\partial z} dz = \frac{u_*}{k} \int_{z_0}^z \frac{dz}{z}$$

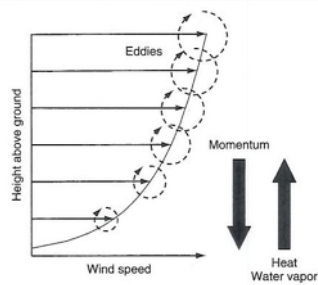
$$\bar{u}(z) = \frac{u_*}{k} \ln \frac{z}{z_0}$$



- Mean wind speed in the surface layer is **decelerated by friction** whose influence is felt aloft through eddy momentum flux
- Varies **logarithmically with height**
- Y-intercept of log-linear plot of SL wind vs  $z$  is  $z_0$ , which we define as the “**roughness length**”

y-intercept of log-linear plot of SL wind vs  $z$  is  $z_0$ , which we define as the “roughness length”

## Roughness Length



Surface	Roughness length (m)
Soil	0.001–0.01
Grass	
Short	0.003–0.01
Tall	0.04–0.10
Crop	0.04–0.20
Forest	1.0–6.0
Suburban	
Low density	0.4–1.2
High density	0.8–1.8
Urban	
Short building	1.5–2.5
Tall building	2.5–10

For neutral conditions

$$(z/u_*) \partial \bar{u} / \partial z = 1/k$$

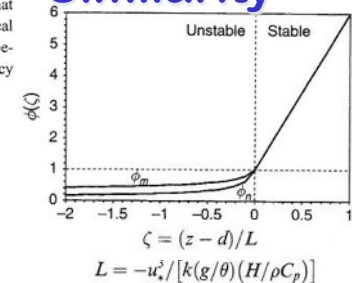
## Monin-Obukhov Similarity

The Monin-Obukhov similarity theory states that when scaled appropriately the dimensionless mean vertical gradients of wind ( $\bar{u}$ ), potential temperature ( $\bar{\theta}$ ), and specific humidity ( $\bar{q}$ ) are unique functions of a buoyancy parameter ( $\zeta$ ):

$$\left[ \frac{k(z-d)}{u_*} \right] \frac{\partial \bar{u}}{\partial z} = \phi_m(\zeta)$$

$$\left[ \frac{k(z-d)}{\theta_*} \right] \frac{\partial \bar{\theta}}{\partial z} = \phi_h(\zeta)$$

$$\left[ \frac{k(z-d)}{q_*} \right] \frac{\partial \bar{q}}{\partial z} = \phi_w(\zeta)$$

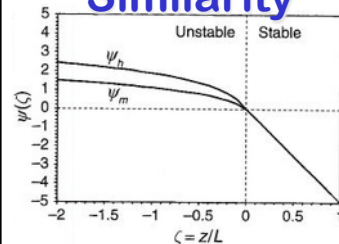


$$\phi_m^2(\zeta) = \phi_h(\zeta) = \phi_w(\zeta) = (1 - 16\zeta)^{-1/2} \quad \text{for } \zeta < 0 \text{ (unstable)}$$

$$\phi_m(\zeta) = \phi_h(\zeta) = \phi_w(\zeta) = 1 + 5\zeta \quad \text{for } \zeta \geq 0 \text{ (stable)}$$

- Empirical adjustment of log-wind profiles to account for buoyancy fluxes (anisotropy)

## Surface-Layer Similarity



$$\bar{u}(z) = \frac{u_*}{k} \left[ \ln \left( \frac{z-d}{z_{0M}} \right) - \psi_m(\zeta) \right]$$

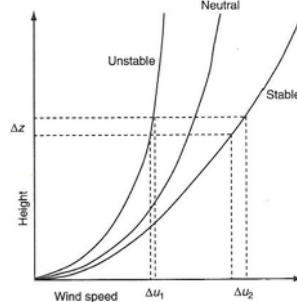
$$\bar{\theta}(z) - \bar{\theta}_s = \frac{\theta_*}{k} \left[ \ln \left( \frac{z-d}{z_{0H}} \right) - \psi_h(\zeta) \right]$$

$$\bar{q}(z) - \bar{q}_s = \frac{q_*}{k} \left[ \ln \left( \frac{z-d}{z_{0W}} \right) - \psi_w(\zeta) \right]$$

Generalized (empirical) vertical profiles

unstable conditions ( $\zeta < 0$ ),  
 $\psi_m(\zeta) = 2 \ln[(1+x)/2] + \ln[(1+x^2)/2] - 2 \tan^{-1} x + \pi/2$   
 $\psi_h(\zeta) = \psi_w(\zeta) = 2 \ln[(1+x^2)/2]$   
 where  $x = (1 - 16\zeta)^{1/4}$

For a stable atmosphere ( $\zeta \geq 0$ ),  
 $\psi_m(\zeta) = \psi_h(\zeta) = \psi_w(\zeta) = -5\zeta$



## Estimation of Turbulent Fluxes

where

$$\tau = \rho K_M (\partial \bar{u} / \partial z)$$

$$H = -\rho C_p K_H (\partial \bar{\theta} / \partial z)$$

$$E = -\rho K_W (\partial \bar{q} / \partial z)$$

$$K_M = k u_* (z-d) / \phi_m(\zeta)$$

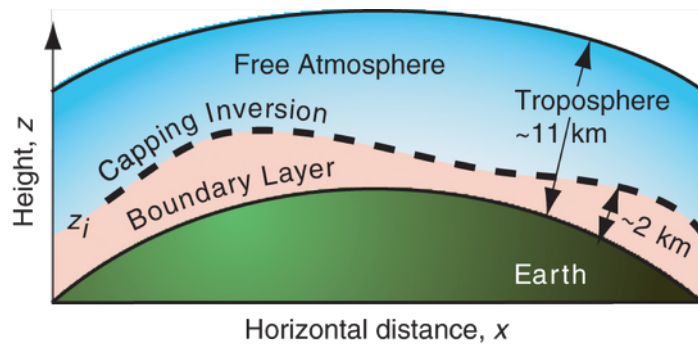
$$K_H = k u_* (z-d) / \phi_h(\zeta)$$

$$K_W = k u_* (z-d) / \phi_w(\zeta)$$

- Fluxes are **driven by gradients** in  $u$ ,  $T$ , and  $q$
- Fluxes are **proportional to friction velocity**
- These are simply **definitions** of  $K_M$ ,  $K_H$ ,  $K_W$
- **Ohm's Law combined with Similarity**

## The Atmospheric Boundary Layer

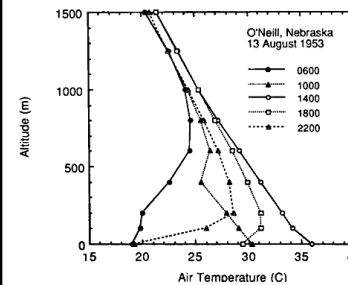
(a.k.a ABL, PBL, CBL, NBL, SBL ...)



*"The layer of atmosphere in turbulent contact with the surface"*

## PBL Temperatures

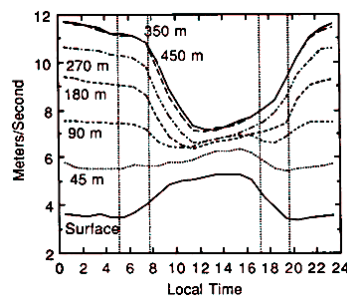
Diurnal Cycle



- **Morning inversion** broken by surface heating
- Shallow ML by 10 AM under RL from **yesterday**
- **Superadiabatic** surface layer at 2 PM
- New **inversion forms near surface** by 6 PM
- Nocturnal BL grows "from the **bottom up**"

## PBL Wind Speeds

Annual Mean Diurnal Cycles



- Surface winds are maximum at midday
- Winds aloft are maximum at night (decoupling)
- Momentum mixing during daytime allows surface friction to be "felt" throughout ML

## Typical Diurnal Cycle of PBL Over Land

